Chapter 3

Radiation



Units of Chapter 3

Types of radiation

Waves

Waves in What?

The Wave Nature of Radiation

The Electromagnetic Spectrum

Thermal Radiation

The Kelvin Temperature Scale

More about the Radiation Laws

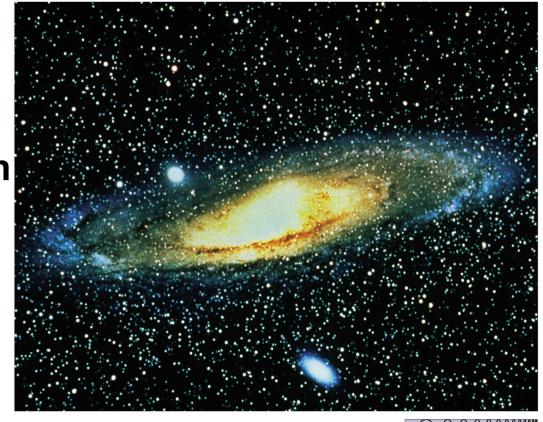
The Doppler Effect

Types of Radiation

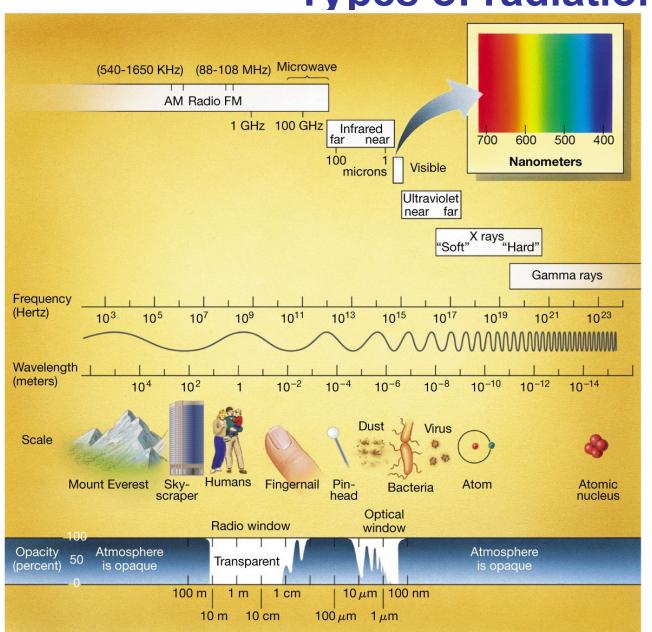
Electromagnetic Radiation: energy transmitted through space as varying electric and magnetic fields

Light, x-rays radio waves, infrared

Particulate radiation beta rays (e-), alpha rays (He) [Not covered here!]



Types of radiation



Electromagnetic radiation

Different ranges have different names

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Types of radiation

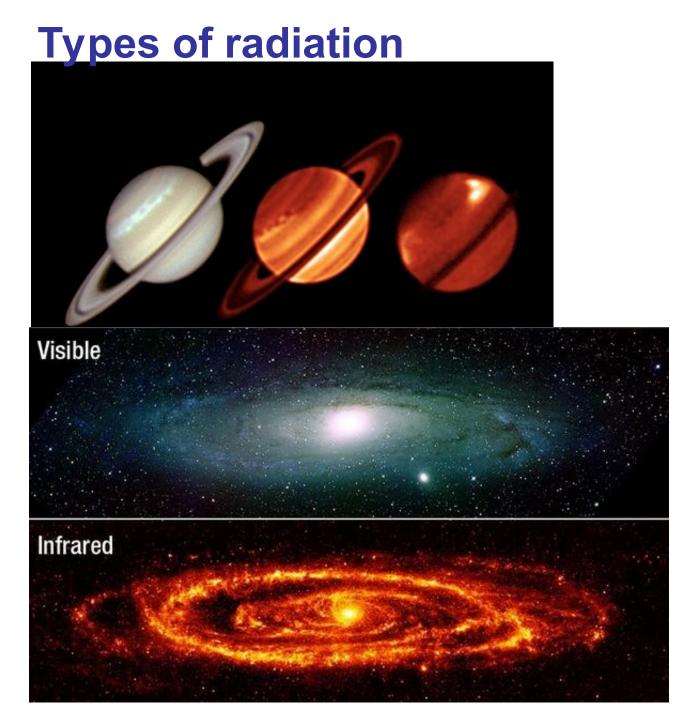
Electromagnetic Radiation may be transmitted, reflected, absorbed, or scattered off of obstructions. How it interacts depends on 1) the type of radiation (radio, IR, etc), 2) the composition.

Visible

Infrared



Astronomical objects in different wavelengths.



Waves

Wave: a travelling disturbance or variation in a medium or field which carries energy.

Types:

Mechanical Electromagnetic Gravitational

sound Light inspiralling BHs

seismic microwaves "chirp"

water x-rays, gamma rays

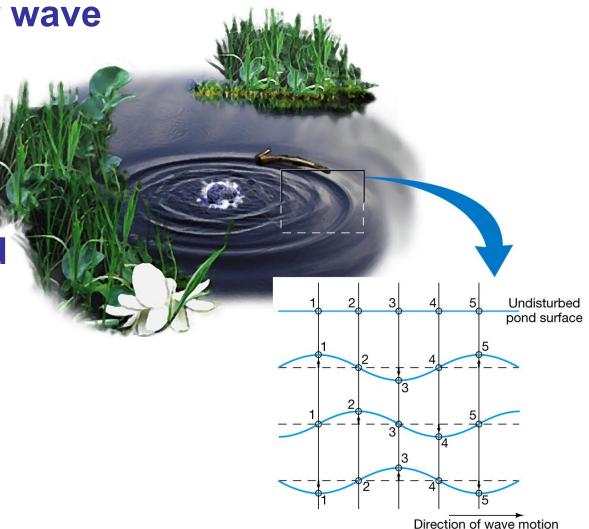
"the wave"

What do they have in common?

Example: water wave

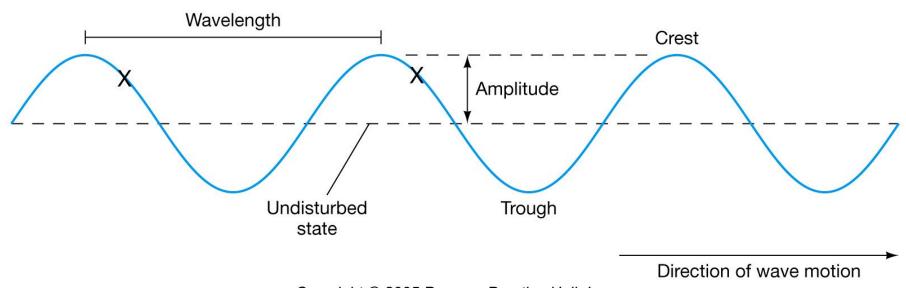
Water just moves up and down

Wave travels and can transmit energy



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Sine waves: waves described by a sine or cosine function. Also called: "sinusoidal"

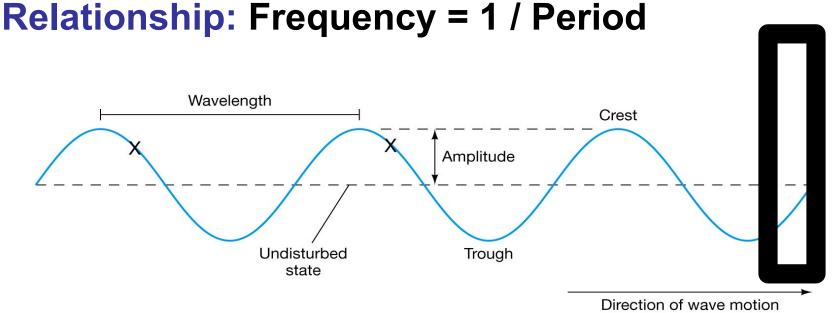


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This graph shows <u>amplitude versus position</u>, but <u>amplitude versus time</u> is ALSO a sinusoidal graph!

Frequency: number of wave crests that pass a given point per second

Period: time between passage of successive crests

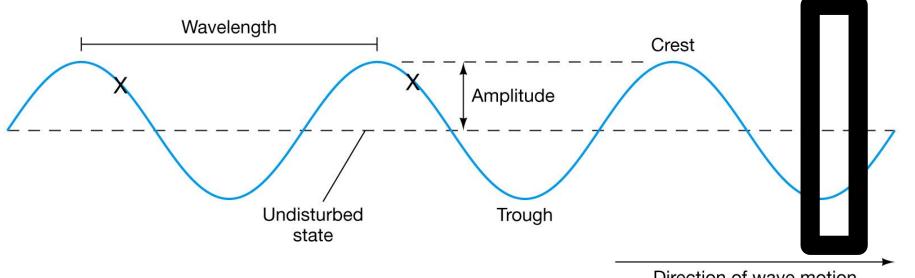


Wavelength: distance between successive crests

Velocity: speed at which crests move

Velocity = Wavelength/Period

Velocity = Wavelength * frequency

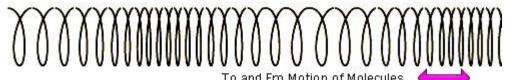


Direction of wave motion

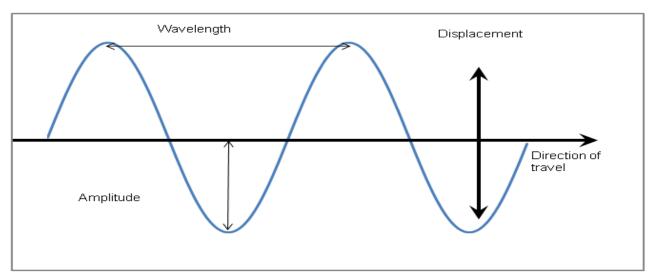
Golden

Rule! $V = \lambda f$

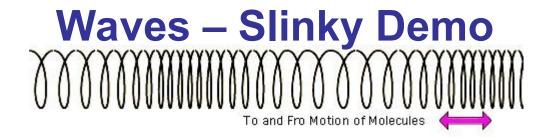
Longitudinal wave: propogates in a direction parallel to the displacement of the medium



Transverse wave: propogates in a direction perpendicular (or transverse) to the displacement of the medium



DEMO: long. and transv. waves in a SLINKY! Standing waves



- 1. Longitudinal pulses
- 2. Transverse pulses note reflected pulse
- 3. Speed of pulse increases with tension
- 4. Superposition of pulses
- 5. Standing waves (superposition with reflected waves)
- 6. Harmonics (N=number of antinodes)
- 7. Polarization

Waves in What?

Water waves, sound waves, and so on, travel in a medium (water, air, ...)

Electromagnetic waves need no medium

Created by accelerating charged particles:

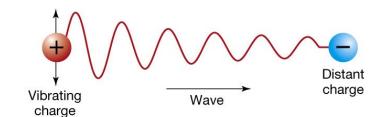
Electric field lines

Field line

Distant charge

Stationary charge

Demo: spark makes radio waves!

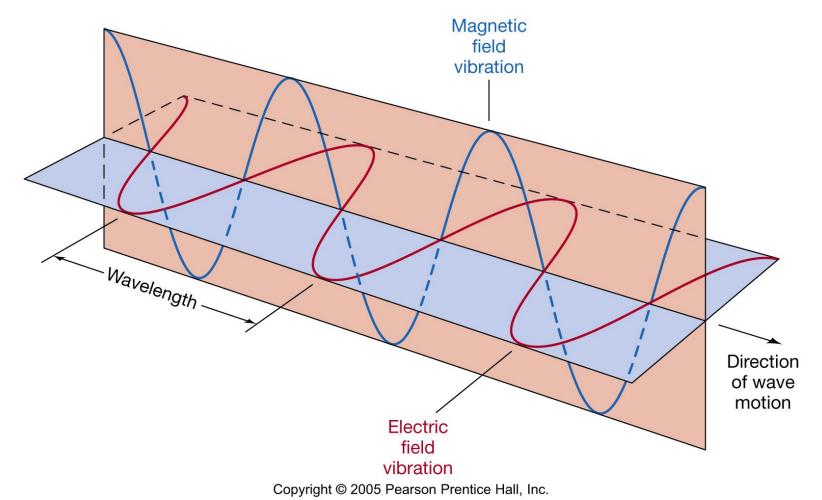


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(c)

Waves in What?

Electromagnetic waves: Oscillating electric and magnetic fields. Changing electric field creates magnetic field, and vice versa



Waves in What?

What is the wave speed of electromagnetic waves?

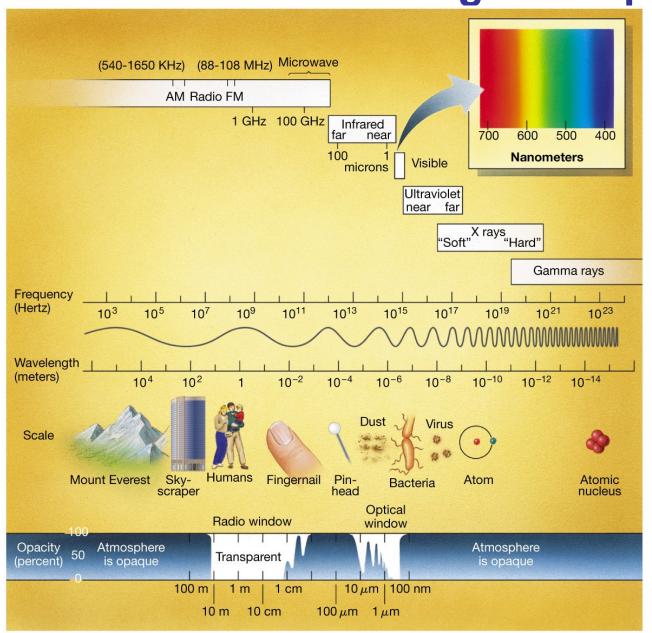
$$c = 3.0 \times 10^8 \text{ m/s}$$

This speed is very large, but still finite; it can take light millions or even billions of years to traverse astronomical distances.

Why special?

- 1) Nature's speed limit.
- 2) A beam of light appears to move at the same speed through a vacuum to any observer.

The Electromagnetic Spectrum



No upper limit on wavelength

High frequency radiation has small wavelength.

High opacity means low transparency.

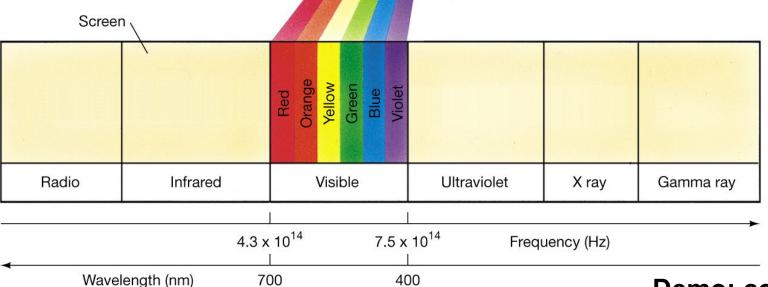
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Electromagnetic spectrum

Refraction: the bending of light at an interface between media.

Dispersion: spreading apart of light into colors.

Visible spectrum:



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Demo: color pinwheel

Slit

Prism

White light

Light as wave or particle

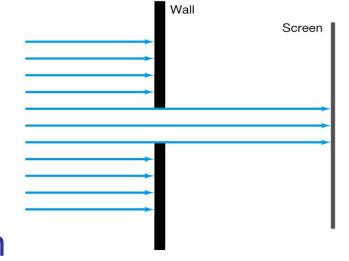
Phenomena best described with waves:

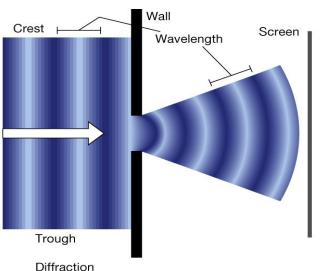
Diffraction = bending of light around corners and slits.

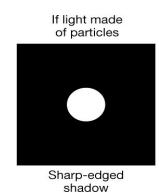
Top: no diffraction

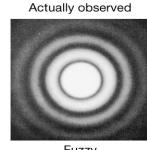
Bottom: diffraction

DEMO: laser diffracted by edges.









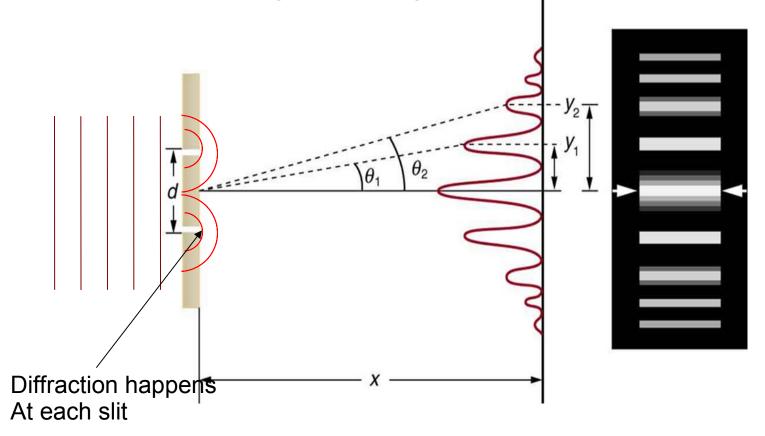
Fuzzy shadow

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Light as wave or particle

Phenomena best described with waves

Interference = two or more waves can combine destructively as well as constructively when they meet at a point.



Light as wave or particle

An example of a phenomenon which is best described with the <u>particle model</u> is ...

The Photoelectric Effect

* Light with a freq above some limit can dislodge efrom the surface of a metal. Just below that limit, no e- dislodged even if the intensity of the light is great!

* Conclusion: light comes in particles called photons with Ephot = hf. (h=6.626x10⁻³⁴ Js)

See [phet.colorado.edu/en/simulation/photoelectric]

Light as wave or particle Phenomena which could be described with the particle and wave models are ...

Reflection

* the bouncing of photons or waves off of a shiny surface such that ...

* angle of incidence = angle of reflection

Refraction (wave model is preferred)

* the slowing and bending of light when travelling not from one medium to another

 $n_1 < n_2$

* Snell's law: $n_1 \sin \Theta_1 = n_2 \sin \Theta_2$

Thermal radiation: the light produced (not reflected) by real objects which depends on the object's temperature and emissivity.
--> Closely approximates blackbody radiation.

Blackbody: absorbs 100% of incident light, and emits light with a blackbody spectrum (continuous with single peak).

Coal is a good approximation of a black body.





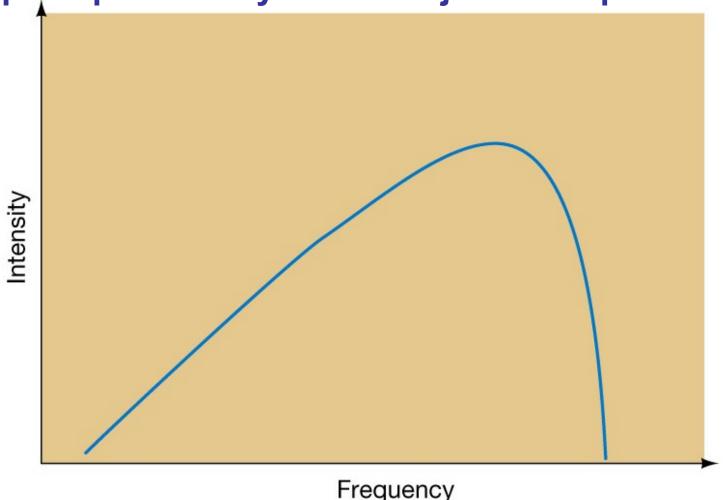
See also "Vantablack" online.

"Vantablack" is a better approximation of a black body (99.96% of light absorbed).





Blackbody Spectrum: radiation emitted by a blackbody, or perfect absorber. The spectrum's shape depends only on the object's temperature.

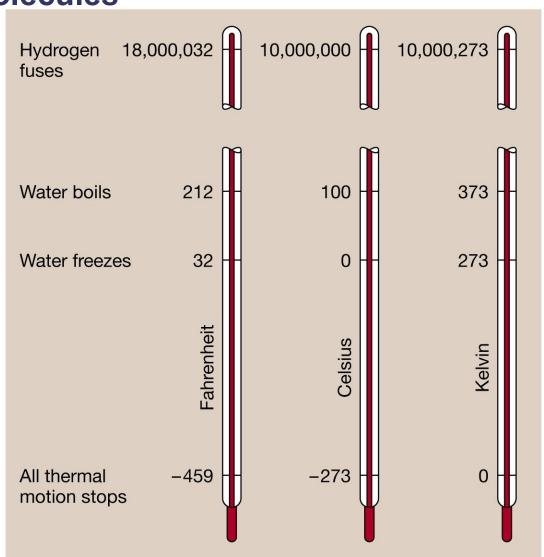


Thermal Radiation Review: Temperature

Temperature: a measure of the energy stored in the random motions of atoms and molecules

Kelvin – an absolute temperature scale:

- All thermal motion ceases at 0 K
- Water freezes at 273
 K and boils at 373

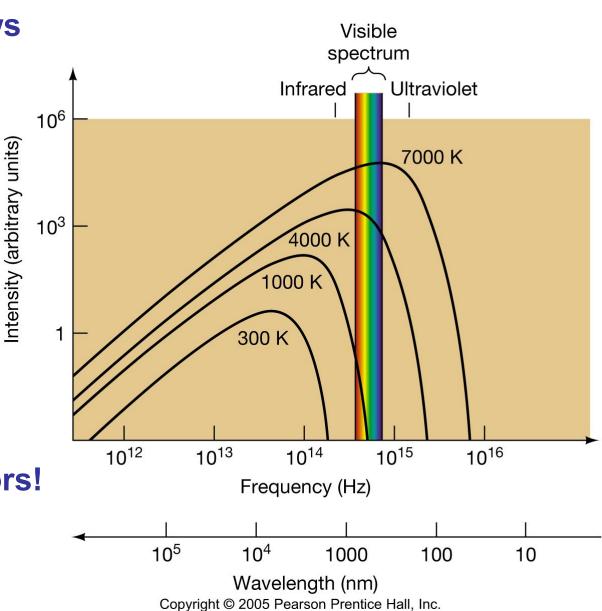


Thermal Radiation Laws

1. Wien's Law:
Peak wavelength is
inversely proportional
to temperature.

$$\lambda_{\text{max}} \sim 1/T$$

This gives us a way to estimate temperatures of stars from their colors!

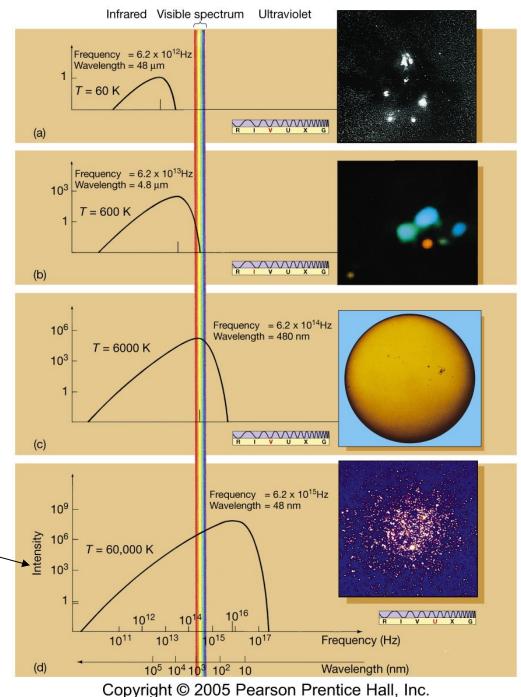


Radiation Laws

2. Stefan's Law: **Total energy emitted** is proportional to the fourth power of temperature; $I \propto T^4$.

Note: intensity scale of curves is logarithmic!

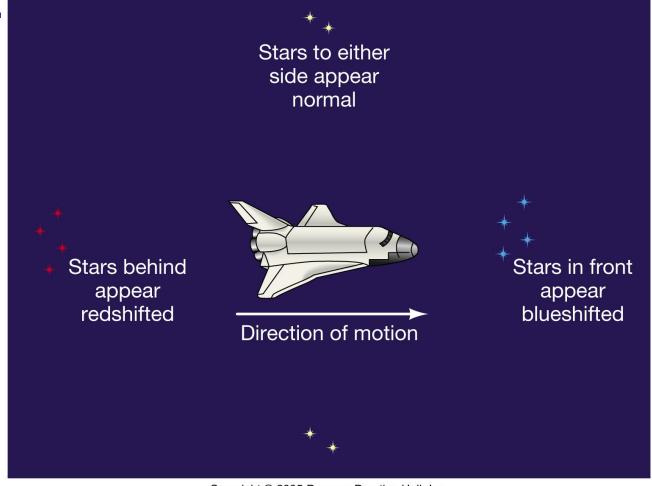
DEMO: lightbulb filament with varying current



3.5 The Doppler Effect

If one is moving toward a source of waves, the wavelengths seem shorter; if moving away, they

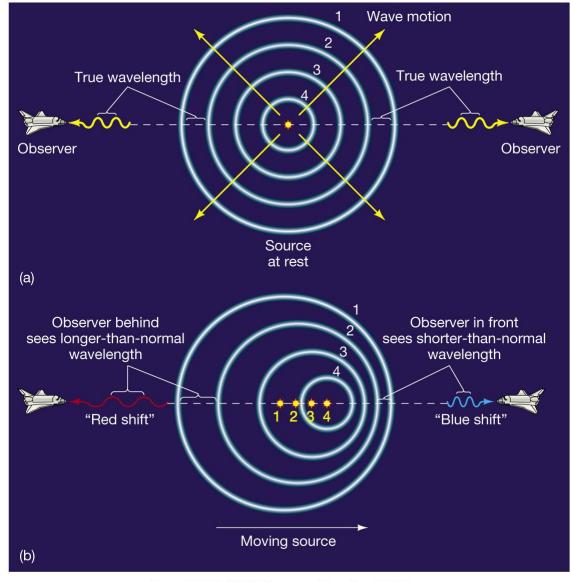
seem longer.



3.5 The Doppler Effect

Depends only on the relative motion of source

and observer:



Summary of Chapter 3

- Wave: period, wavelength, amplitude
- Electromagnetic waves created by accelerating charges
- Visible spectrum is different wavelengths of light
- Entire electromagnetic spectrum:

radio waves, infrared, visible light, ultraviolet, X rays, gamma rays

Summary of Chapter 3, cont.

- Can tell the temperature of an object by measuring its thermal radiation
- Doppler effect can change perceived frequency of radiation
- Doppler effect depends on relative speed of source and observer