## Physics 1051. Planetary Astronomy Quiz 4 REVIEW.

## Light and spectroscopy

| 1. | Which of these properties for waves on a string are measured in length units (e.g., meters, mm, etc)? (Circle more than one.)    |  |  |  |  |  |  |  |  |  |
|----|--|--|--|--|--|--|--|--|--|--|
|    | (a) speed (b) frequency (c) amplitude (d) wavelength (e) polarization  |  |  |  |  |  |  |  |  |  |
| 2. | Which type of electromagnetic radiation has the longest wavelength?  |  |  |  |  |  |  |  |  |  |
|    | (a) Radio (b) visible (c) ultraviolet (d) infrared (e) gamma rays  |  |  |  |  |  |  |  |  |  |
| 3. | What are the colors contained in white light, ordered from low to high frequency? (First letter only. Give at least six colors.) |  |  |  |  |  |  |  |  |  |
| 4. | Which type of electromagnetic radiation has the lowest frequency?  |  |  |  |  |  |  |  |  |  |
|    | (a) Radio (b) visible (c) ultraviolet (d) infrared (e) gamma rays  |  |  |  |  |  |  |  |  |  |
| 5. | Which type of electromagnetic radiation has the highest frequency?   |  |  |  |  |  |  |  |  |  |
|    | (a) Radio (b) visible (c) ultraviolet (d) infrared (e) gamma rays  |  |  |  |  |  |  |  |  |  |
| 6. | Name a behavior of light which is characteristic of waves but not particles.   |  |  |  |  |  |  |  |  |  |
| 7. | Which of these is a type of electromagnetic radiation which is harmful to living tissues? (Circle one)                           |  |  |  |  |  |  |  |  |  |
|    | (a) gamma rays b) infrared c) radio d) visible light e) cosmic rays  |  |  |  |  |  |  |  |  |  |
| 8. | The temperature of an ideal blackbody can be measured from the its spectrum.   |  |  |  |  |  |  |  |  |  |
|    | (a) emission lines in  |  |  |  |  |  |  |  |  |  |
|    | (b) absorption lines in  |  |  |  |  |  |  |  |  |  |
|    | (c) peak intensity of  |  |  |  |  |  |  |  |  |  |
|    | <ul><li>(d) wavelength of peak intensity for</li><li>(e) width of absorption lines in</li></ul>                                  |  |  |  |  |  |  |  |  |  |
| 9. | The pitch of a train horn will drop as the train passes because of the effect.   |  |  |  |  |  |  |  |  |  |

| 10. | The spectrum of a star moving away from us will be compared to a similar star at rest.   |
|-----|--|
|     | <ul><li>(a) redshifted</li><li>(b) blueshifted</li></ul>   |
|     | (c) neither redshifted or blueshifted  |
|     | (d) brighter   |
|     | (e) dimmer   |
| 11. | The spectrum of a star moving toward us will be compared to a similar star at rest.  |
|     | (a) redshifted   |
|     | (b) blueshifted  |
|     | (c) neither redshifted or blueshifted  |
|     | (d) brighter   |
|     | (e) dimmer   |
| 12. | The amount of radiation leaving a blackbody depends on temperature raised to the power.  |
|     | (a) 1st (b) 2nd (c) 3rd (d) 4th (e) 5th  |
| 13. | Varying the current flowing through a lightbulb will change the filaments temperature and allow one to demonstrate                                       |
|     | (a) Wien's law (b) Stefan's Law (c) Kirchoff's laws (d) Wien's law and Stefan's law (e) Newton's law   |
| 1   | Light and telescopes   |
| (NO | ΓE: Only those with ** are important.)   |
|     |  |
| 1.  | The diameter of the pupil is about 0.2 inches. How many times fainter than the naked-eye limit can be seen with a 1-inch diameter telescope? (Hint: LGP) |
|     | (a) 2 times (b) 4 times (c) 5 times (d) 25 times   |
| 2.  | The Light Gathering Power (LGP) of a telescope of mirror diameter D is proportional to   |
|     | (a) D (b) $D^2$ (c) $1/D$ (d) the eye's pupil diameter   |
| 3.  | The following are designs of reflecting telescopes except the  |

|              | (a)     | refracto              | or (    | (b) New   | tonian     | (c) Pr                                | ime focu  | us (d)    | ) Casse  | grain     | (e) C      | oude                 |       |
|--------------|---------|-----------------------|---------|-----------|------------|---------------------------------------|-----------|-----------|----------|-----------|------------|----------------------|-------|
| 4.           |         | is the er than        |         |           |            | can see                               | using a   | telescope | e with   | a lens o  | diamete    | r 10 ti              | imes  |
|              | (a)     | 1.0                   | (b) -6. | 0 (       | c) 4.0     | (d) 6.0                               | (e)       | ) 11.0    |          |           |            |                      |       |
| 5.           | The v   | various t             | ypes of | freflecti | ng telesc  | opes can                              | be distin | nguished  | by the   | position  | n of thei  | r                    | ·     |
|              | \ /     | focal pla<br>clock dr |         | (b) pri   | mary mi    | rror                                  | (c) secor | ndary mi  | rror     | (d) fin   | der scop   | oes                  | (e)   |
| 6.           | The r   | esolutio              | on of a | telescop  | e impro    | ves as yo                             | u increa  | se        | <u> </u> |           |            |                      |       |
|              | . ,     | the mag<br>double s   |         |           | . ,        | e seconda<br>length                   | -         |           |          | ` '       |            | nce to               | the   |
| 7.           | ** In   |                       | (       | lesign o  | f reflecti | ing telesc                            | cope, the | ere is a  | whole t  | hrough    | the cer    | nter of              | the   |
|              | (a)     | Cassegra              | ain     | (b) ref   | fractor    | (c) P                                 | rime      | (d) Cou   | ıdé      | (e) Ne    | wtonian    | L                    |       |
| 8.           | ** Di   | ffraction             | n and i | refractio | n are ex   | amples o                              | f         | _         |          |           |            |                      |       |
|              | ` '     | _                     | _       |           | `          | tricks (e) energ                      |           | re (c)    | wave j   | propert   | ies of lig | $\operatorname{ght}$ | (d)   |
| 9.           | Light   | comes                 | in disc | rete 'pie | eces' of e | nergy ca                              | lled      |           |          |           |            |                      |       |
|              | (a) .   | Joules                | (b)     | keppers   | (c)        | Planck's                              | (d) 1     | photons   | (e)      | bullets   |            |                      |       |
| 10.          |         | esides vi<br>n's atmo |         | _         | ~ -        | electron                              | nagnetic  | radiation | n which  | n is leas | t absorb   | oed by               | the   |
|              | (a) 1   | radio wa              | aves    | (b) X-    | rays       | (c) gamr                              | na-rays   | (d) c     | osmic-r  | ays       | (e) infr   | ared w               | aves  |
| 11.          | ** W    | hich wa               | ve pro  | perty of  | light is   | often me                              | asured i  | n nanom   | eters a  | nd is re  | elated to  | color                | ?     |
|              | (a)     | frequenc              | ey      | (b) ener  | gy (       | (c) wavele                            | ength     | (d) c, t  | the spe  | ed of lig | ght (      | e) reds              | shift |
| $\mathbf{T}$ | he Su   | ın                    |         |           |            |                                       |           |           |          |           |            |                      |       |
| it is 1      | related | l to assi             | gned re | eading o  | or homev   | t yet been<br>vork ques<br>n's surfac | stions.   |           |          | may be    | asked o    | n a qu               | iz if |

(d) 5800 K

(e) 10,000 K

(c) 3800 K

(a) 600 K

(b) 2000 K

| 2.  | This layer of the Sun's atmosphere includes gas with temperatures ranging from 15,000 K to about 1,000,000 K.          |
|-----|--|
|     | (a) photosphere (b) radiative zone (c) transition region (d) chromosphere (e) convective zone                          |
| 3.  | Above the radiative zone of the Sun is a zone where heat is transferred upward by bulk motion of gas, a process called |
|     | (a) conduction (b) convection (c) radiation (d) projection (e) reflection  |
| 4.  | The Sun generates all of its energy in a region called the   |
|     | (a) core (b) fun zone (c) radiative zone (d) convective zone (e) nucleus   |
| 5.  | The most powerful, short-lived explosions on the Sun's surface are called  |
|     | (a) coronal holes (b) flares (c) prominances (d) filaments (e) fusion  |
| 6.  | Which of these surface features on the Sun is the most short-lived?  |
|     | (a) sunspots (b) flares (c) supergranules (d) prominances (e) coronal holes  |
| 7.  | When a gas is maintaining a stable, spherical shape, gravity is balanced by  |
|     | (a) temperature (b) pressure (c) density (d) frictional forces (e) electrical sources                                  |
| 8.  | Evidence for convection on the Sun is seen in bubble-like features about $1000~\mathrm{km}$ across called              |
|     |  |
|     | (a) flares (b) Texans (c) granules (d) prominences (e) sunspots  |
| 9.  | The inhibition (prevention) of convection in regions of strong magnetic fields gives rise to                           |
|     | (a) sunspots (b) prominences (c) flares (d) granules (e) the sunspot cycle   |
| 10. | The Sun's chromosphere is more difficult to observe (fainter) than the photosphere because it is                       |
|     | (a) more colorful (b) cooler (c) farther away (d) more diffuse (less dense) (e) eclipsed by the Sun                    |
| 11. | What is the deepest layer that we can see of the Sun in visible wavelengths?   |
|     | (a) corona (b) chromosphere (c) photosphere (d) convective zone (e) radiative zone                                     |
| 12. | Which layer of the Sun emits most of the photons that reach our eyes directly?   |
|     | (a) corona (b) chromosphere (c) photosphere (d) convective zone (e) radiative zone                                     |

| 13. | T or F. The density and temperature in the solar corona are much higher than in the photosphere.  |
|-----|---|
| 14. | Name a region of the Sun that produces an emission line spectrum, in accordance with Kirchoff's laws.   |
| 15. | In what surface feature of the Sun would you expect to observe absorption lines that are split into 3 lines because of a strong magnetic field? |
| 16. | The nearest star to the Earth can be easily resolved by telescopes. It is called  |
| 17. | What provides the most direct evidence of nuclear reactions currently occuring in the Sun's core?   |
|     | (a) visible light emitted from the core   |
|     | (b) gamma rays emitted from the core  |
|     | (c) x-rays  |
|     | (d) helioseismic vibrations on the surface  |
|     | (e) neutrinos emitted from the core   |
| 18. | The <i>number</i> of sunspots on the Sun increases and decreases with a period of about   |
| 19. | The latitude of sunspots on the Sun increases and decreases with a period of about  |
| 20. | After one, 11 year sunspot cycle, things are back to the starting state except that the of the sunspot pairs is reversed.                       |
| 21. | The CME's from the Sun can lead to on Earth.  |
|     | (a) coronas (b) annihilation (c) auroras (d) migraines (e) helioseismology  |
| 22. | What is the name of the particular nuclear fusion process that provides most of the Sun's power?  |
| 23. | T or F. Since neutrinos can pass through light years of lead without obstruction, we can't construct a neutrino detector on Earth.              |